

Energy Flux Diagnostic and Predictions

Chris Kruse

Ron Smith

Energy Flux

- Energy flux is the product of perturbation pressure and vertical velocity:

$$EF_z = p'w'$$

- In idealized mountain wave studies, the perturbation is defined relative to some mean value that doesn't vary in space
- In the atmosphere and in realistic simulations, large scale pressure variations exist in geostrophic balance with large scale winds
- Demeaning fields leaves large scale perturbations unrelated to gravity waves

Perturbation Quantities

- To address this, we high pass filter
- Filtering is performed using a 2-D isotropic Gaussian high pass filter in Fourier space and inverting
- With the high passed perturbation field, the low passed field can be found by subtracting the high passed field from the full field
- In this way, fields are partitioned into a smoothly varying background part and the high passed part retaining perturbations from gravity waves:

$$p = p_{lp} + p_{hp} = \bar{p} + p'$$

Filtering Caveats

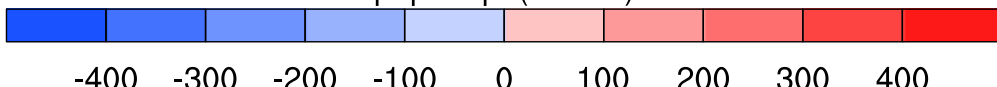
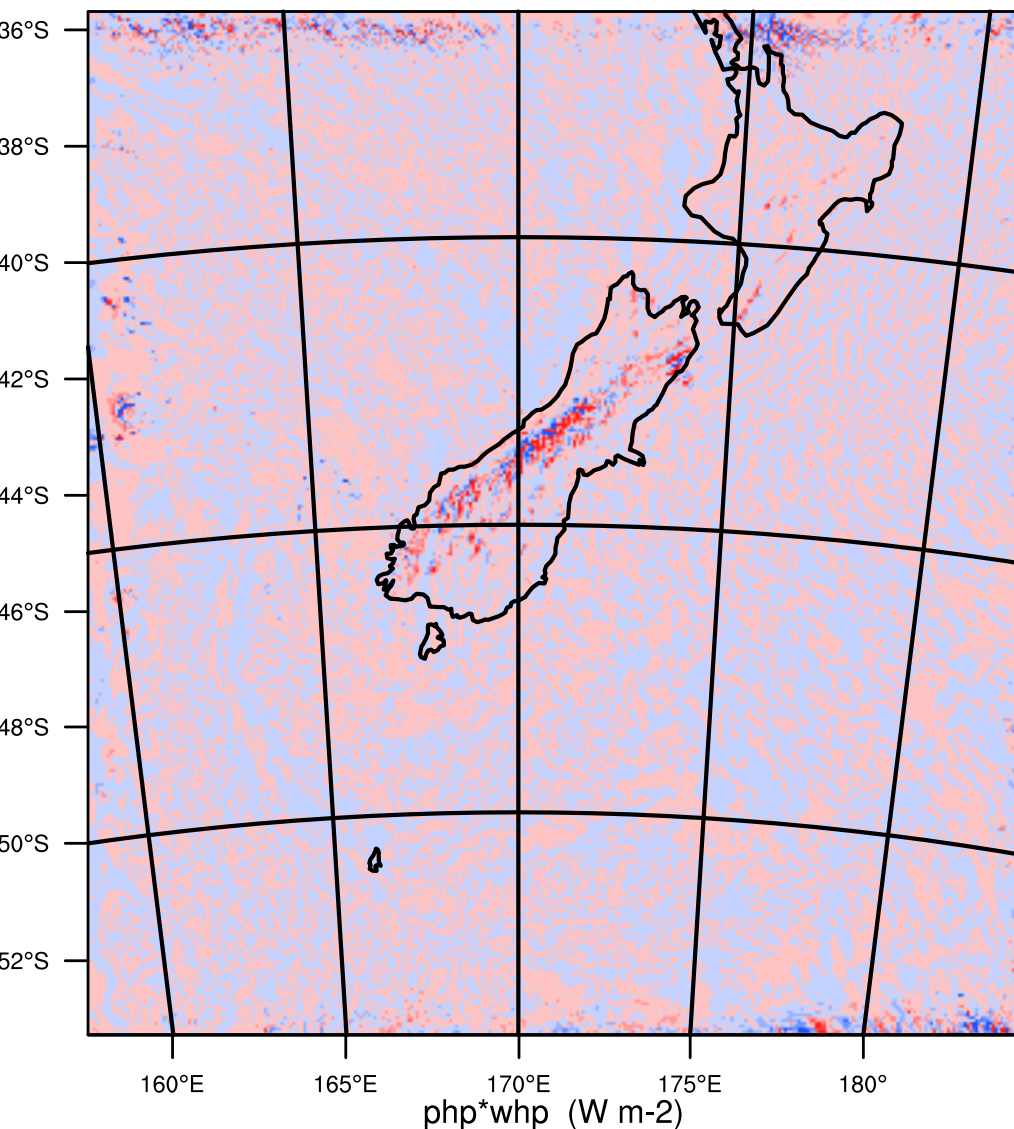
- **Only use one level and one time to separate between large and small scale features**
- **Edge artifacts exist due to aperiodic boundaries**
- **Some sensitivity to cutoff length scale**

EF Calculation

- The energy flux is calculated in four steps:
 1. Model variables interpolated to constant height levels
 2. High pass filtering performed to produce perturbation fields
 3. Perturbation pressure and vertical velocity fields are multiplied point-wise
 4. An additional 150 km low pass filter is applied
 - Evaluates EFz sign locally while retaining spatial extent
 - **Low passed EFz values highly sensitive to LP length scale**

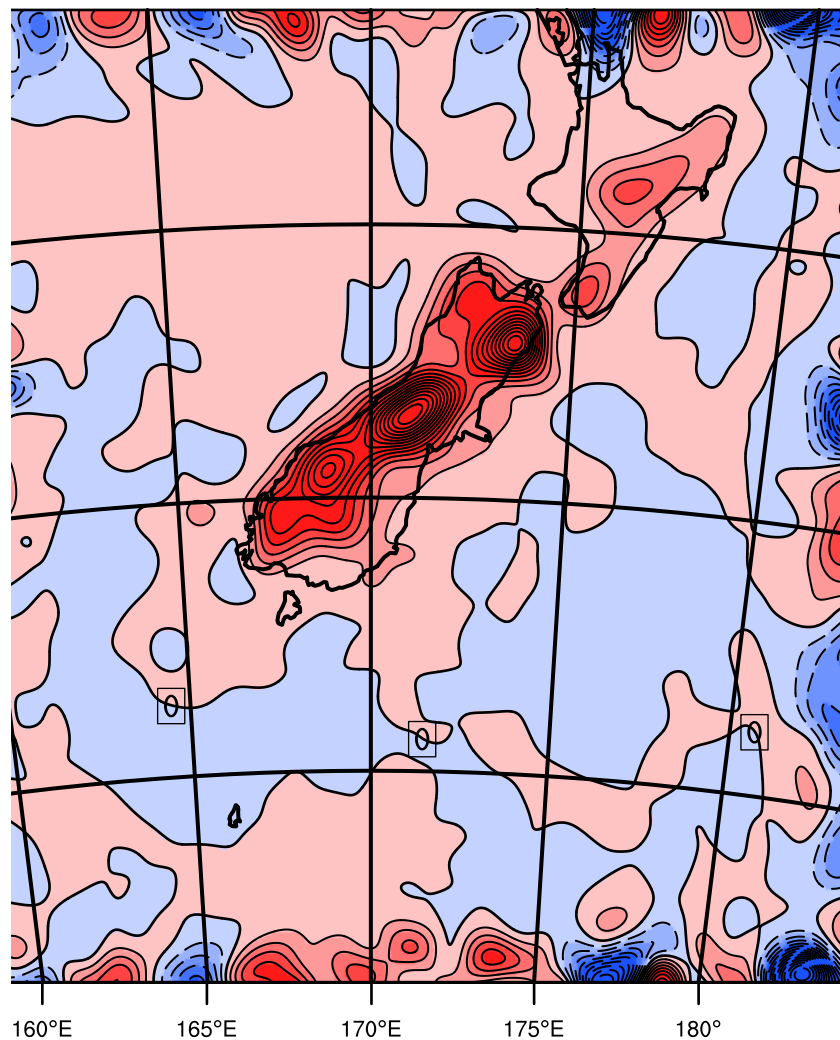
Raw EFz

php*whp (W m-2) at 4 km

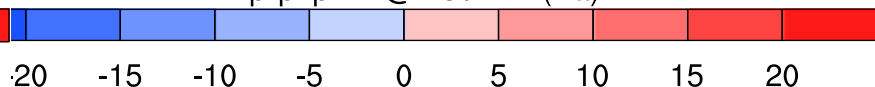


Low Passed EFz Diagnostic

*php LP @ 150 km (Pa) at 4 km
*php LP @ 150 km (Pa) at 4 km



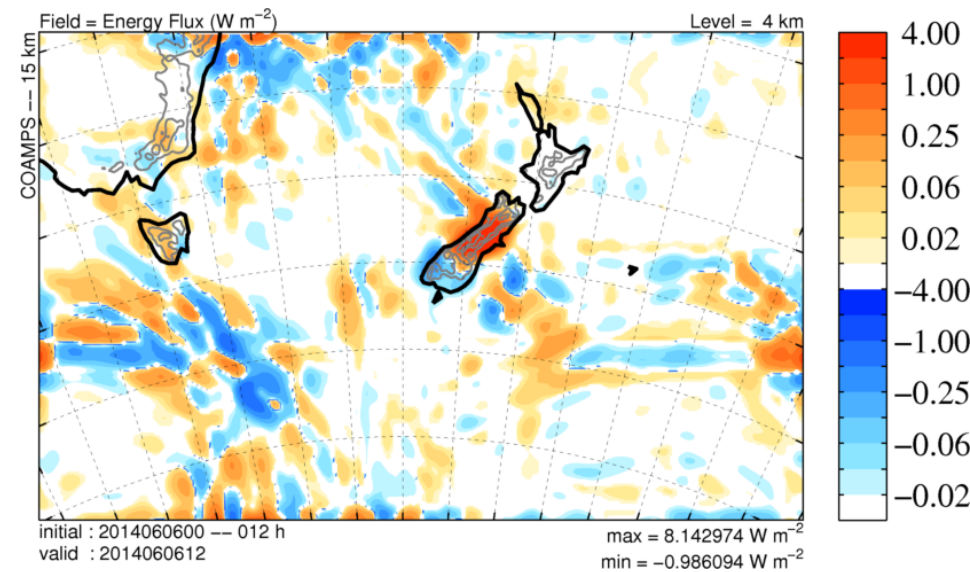
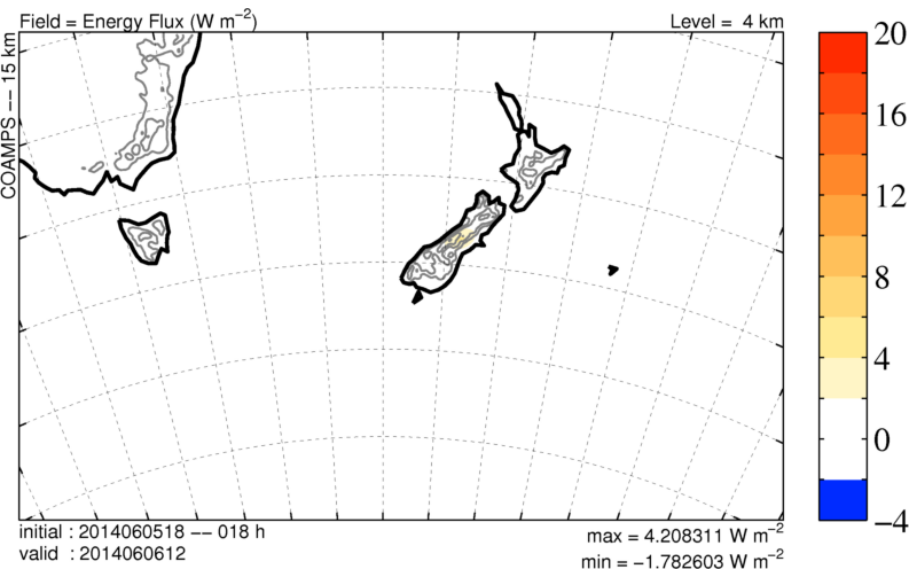
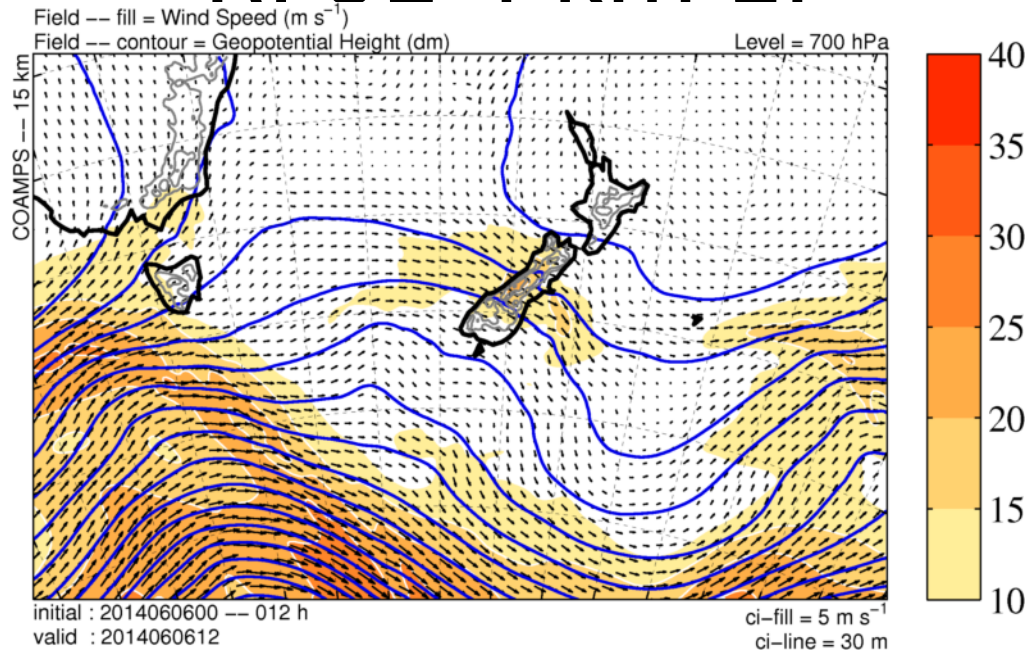
whp*php LP @ 150 km Contours: -85 to 85 by 5
whp*php LP @ 150 km (Pa)



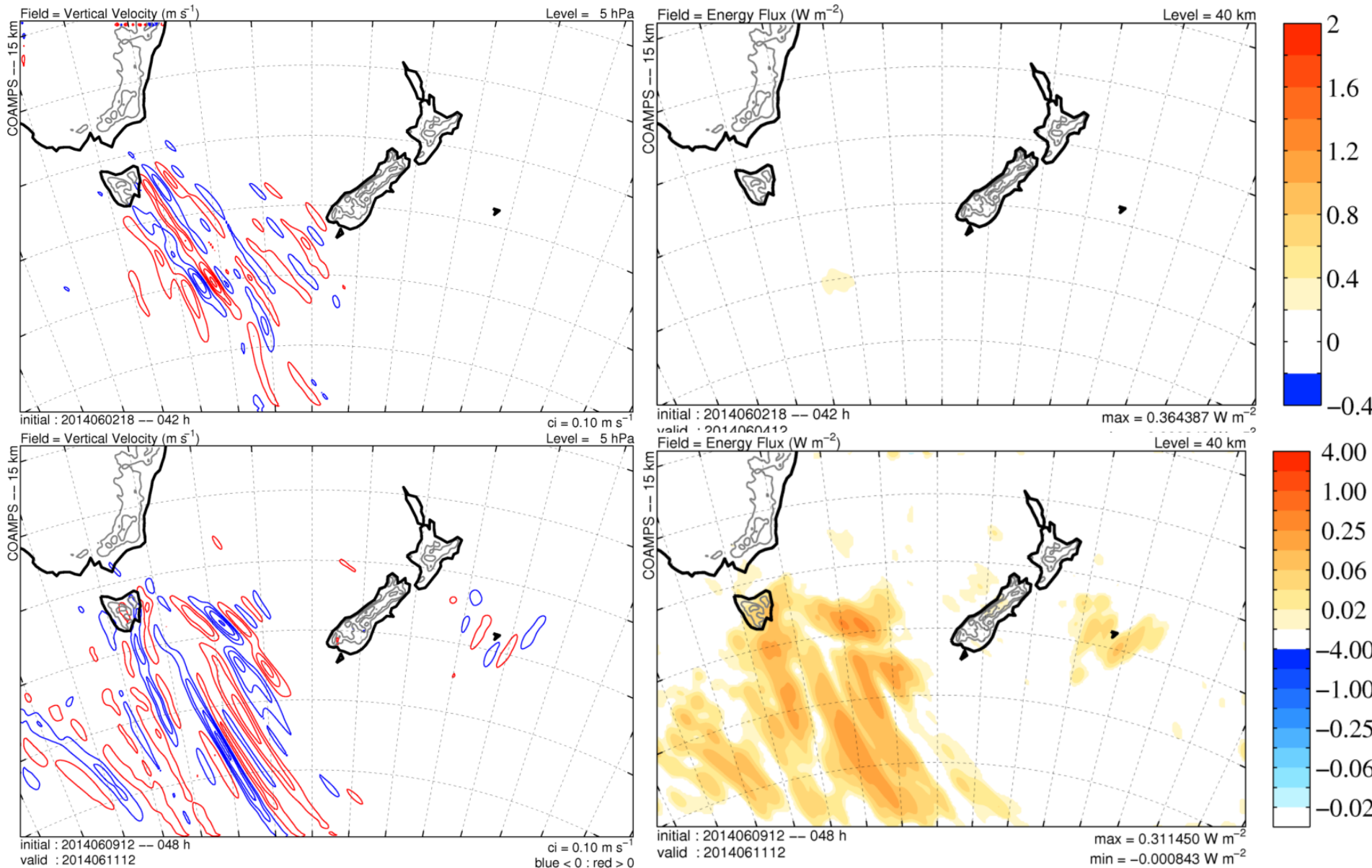
Operational EF Diagnostic Status

- Currently, this diagnostic has been implemented in three operational models:
 1. 15 km COAMPS
 2. 12 km NIWA
 3. 6 km Innsbruck WRF
- Initially, contour spans were targeted at significant orographic cases
- Saw very little flux values with the weaker flow during the beginning of the project
- Logarithmic contour intervals implemented by Alex were quite revealing

RF01 4 km EF



Non-Orographic Examples



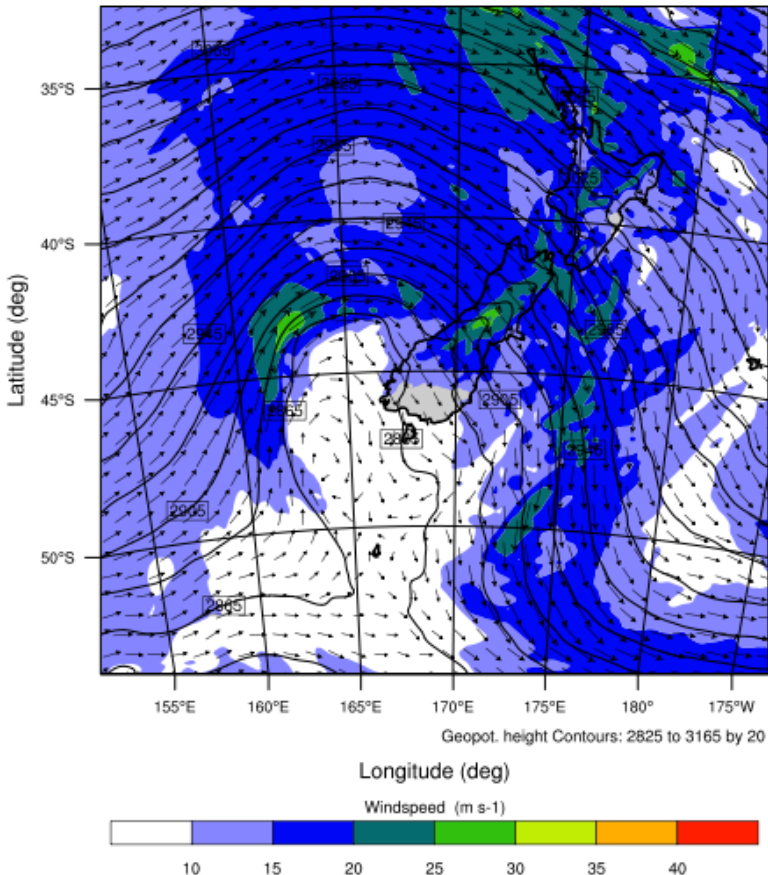
700 hPa Winds Prediction Comparison for RF05

6 km Innsbruck WRF

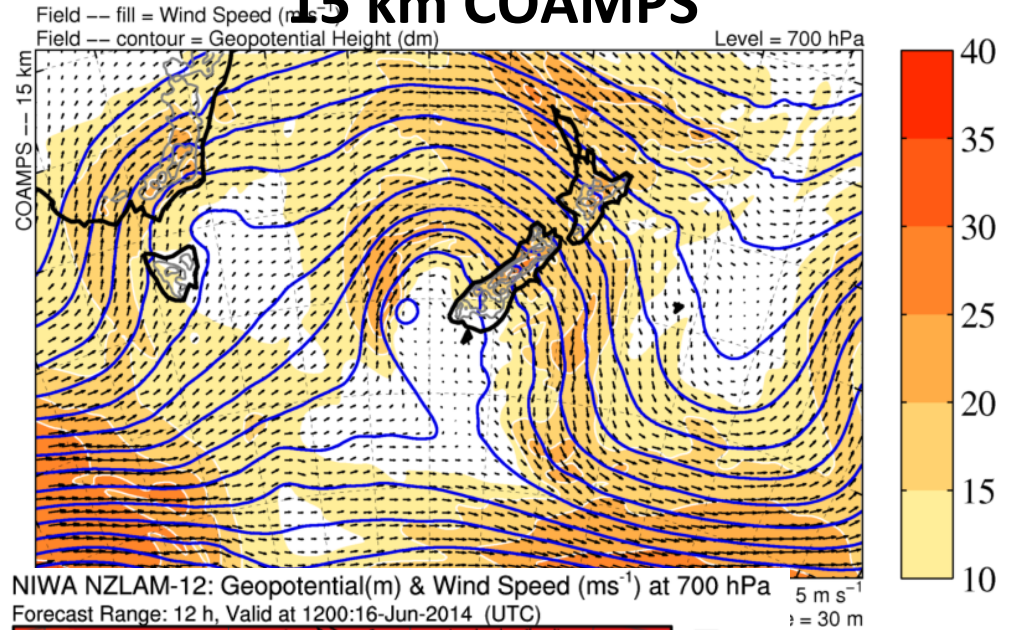
WRF

Init: 2014-06-16_00:00:00
Valid: 2014-06-16_12:00:00

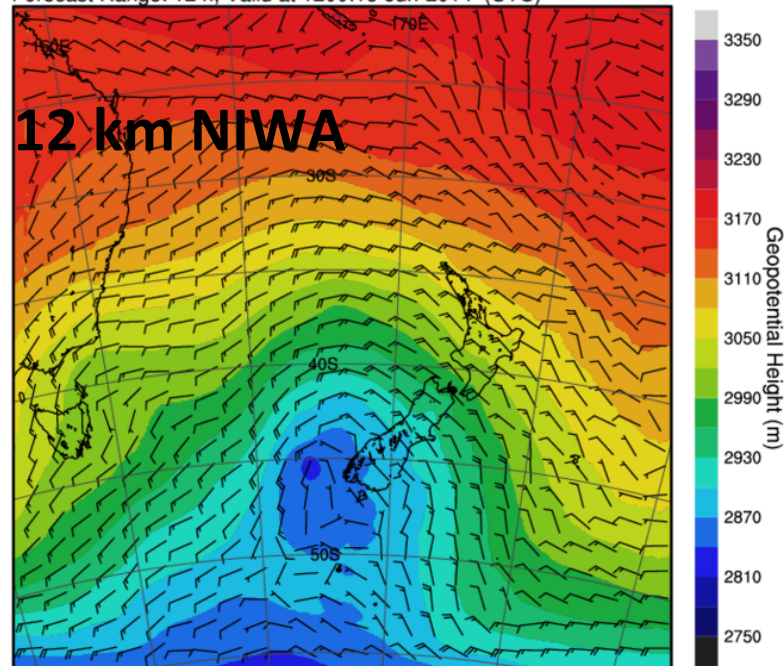
Windspeed (m s⁻¹) at 700 hPa
Geopot. height (m) at 700 hPa
Wind Vectors (m s⁻¹) at 700 hPa



15 km COAMPS



NIWA NZLAM-12: Geopotential(m) & Wind Speed (ms⁻¹) at 700 hPa
Forecast Range: 12 h, Valid at 1200:16-Jun-2014 (UTC)



12 km NIWA

4 km EF Prediction Comparison for RF05

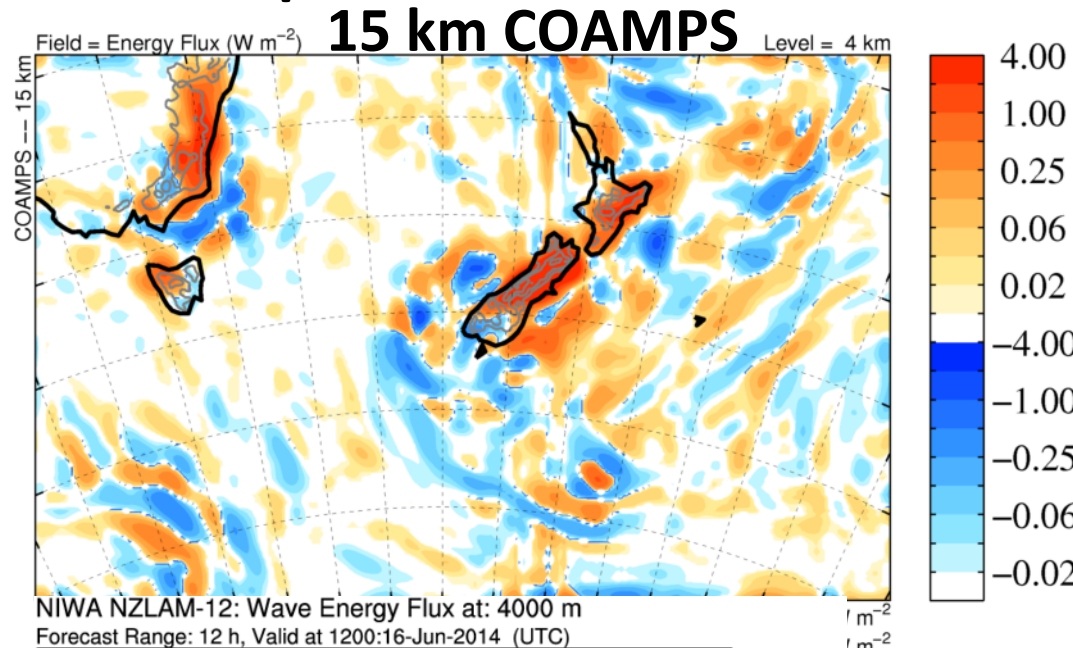
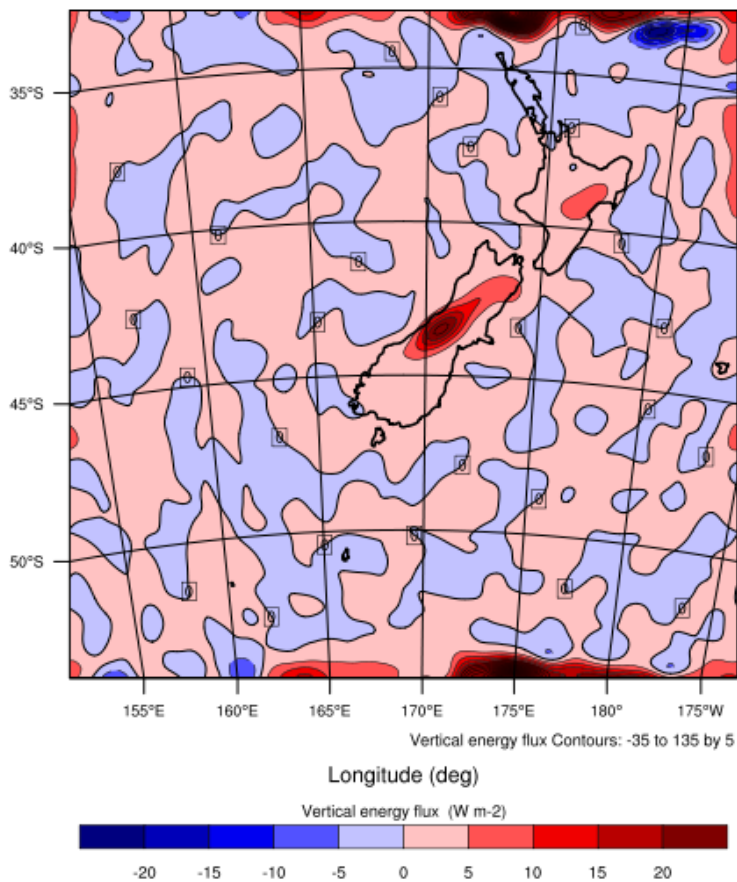
6 km Innsbruck WRF

WRF

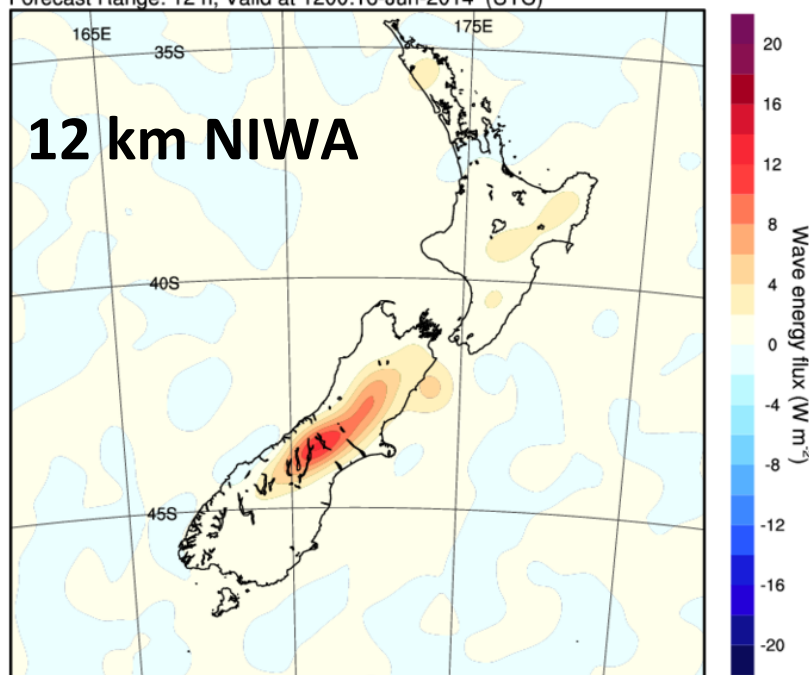
Init: 2014-06-16_00:00:00

Valid: 2014-06-16_12:00:00

Vertical energy flux (W m⁻²) at 4 km
Vertical energy flux (W m⁻²) at 4 km



NIWA NZLAM-12: Wave Energy Flux at: 4000 m
Forecast Range: 12 h, Valid at 1200:16-Jun-2014 (UTC)



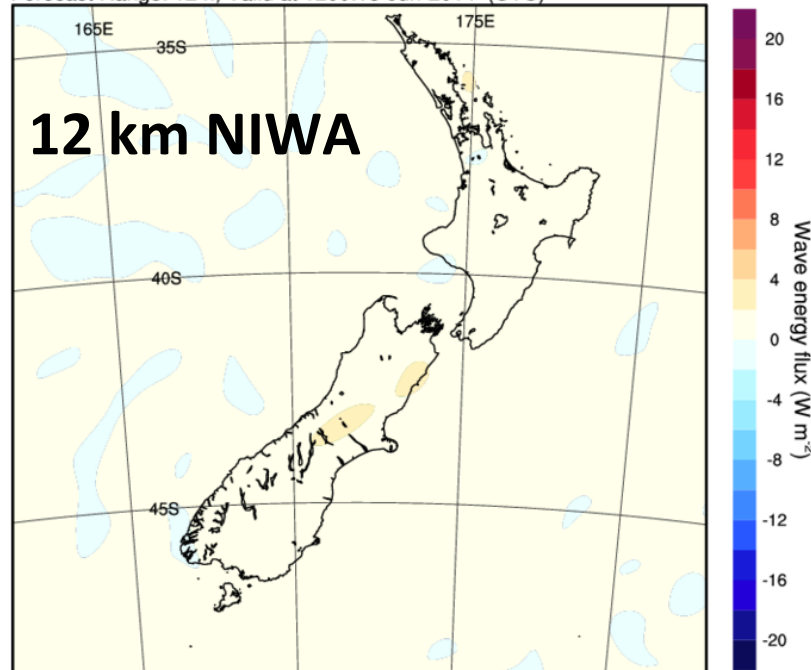
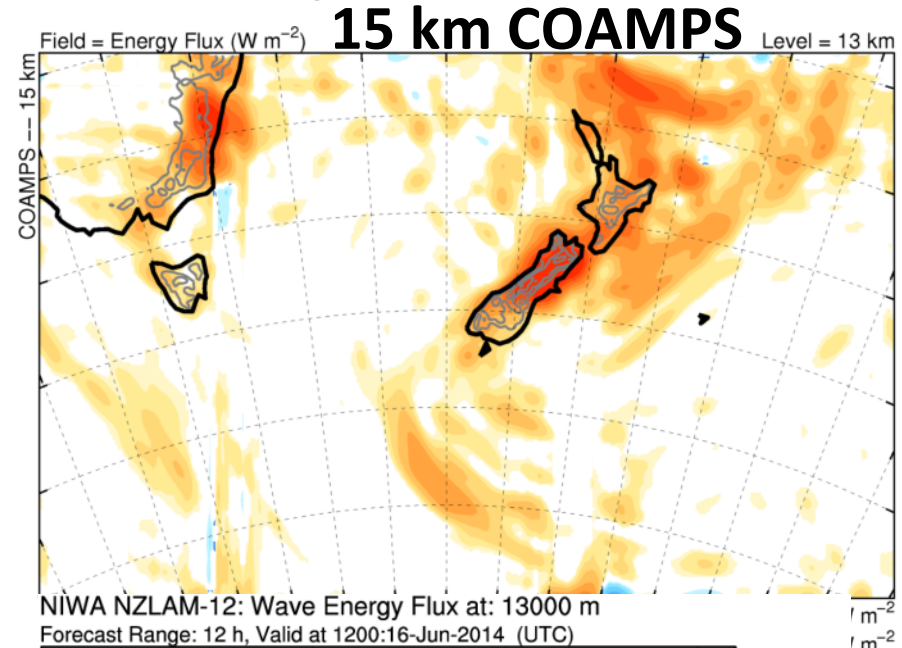
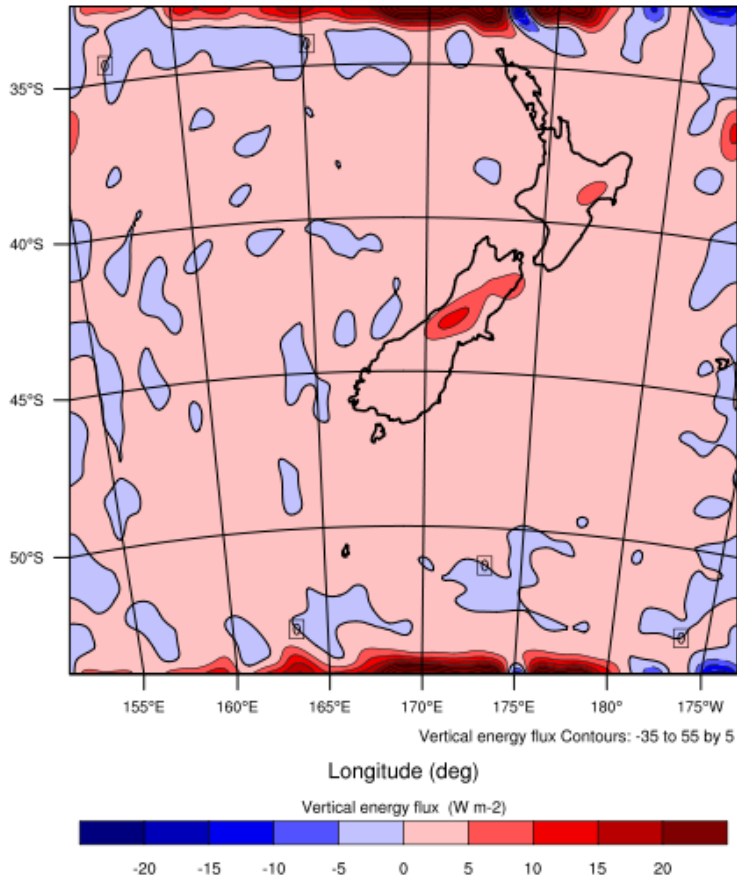
13 km EF Prediction Comparison for RF05

6 km Innsbruck WRF

WRF

Init: 2014-06-16_00:00:00
Valid: 2014-06-16_12:00:00

Vertical energy flux (W m⁻²) at 13 km
Vertical energy flux (W m⁻²) at 13 km



RF05 EF Prediction Comparison

16 June 2014 @ 12 Z, 00Z Initialization

W m-2	4 km EF	13 km EF
COAMPS 15 km	17.8	10.6
NIWA 12 km	12	2
6 km Innsbruck WRF	25	10

16 June 2014 @ 15 Z, 00Z Initialization

W m-2	4 km EF	13 km EF
COAMPS 15 km	18.1	14.4
NIWA 12 km	12	2
6 km Innsbruck WRF	35	10

Conclusions

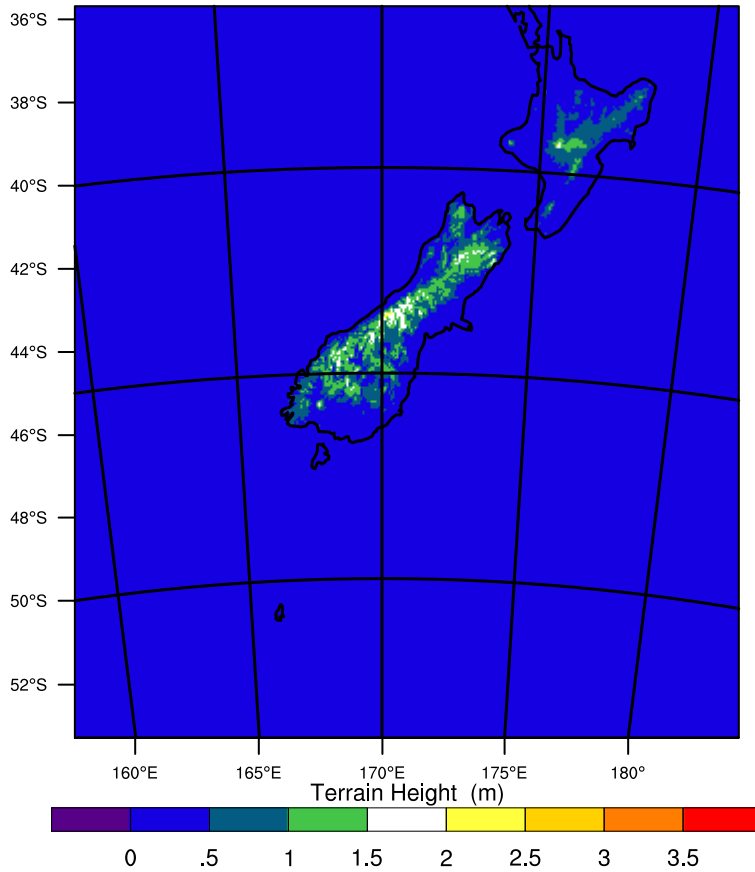
- Strong low level EF result from cross barrier flow
- Strongest fluxes are seen in orographic cases, but non-orographic wave structures do have weaker but positive EF
- EF values are comparable between the NIWA and NRL model
- 6 km resolution simulations produce EF values $>$ lower resolution models
 - Suggests shorter unresolved wavelengths have an important EF contribution
- Interesting weak negative fluxes frequently simulated, but only in the troposphere

Extra

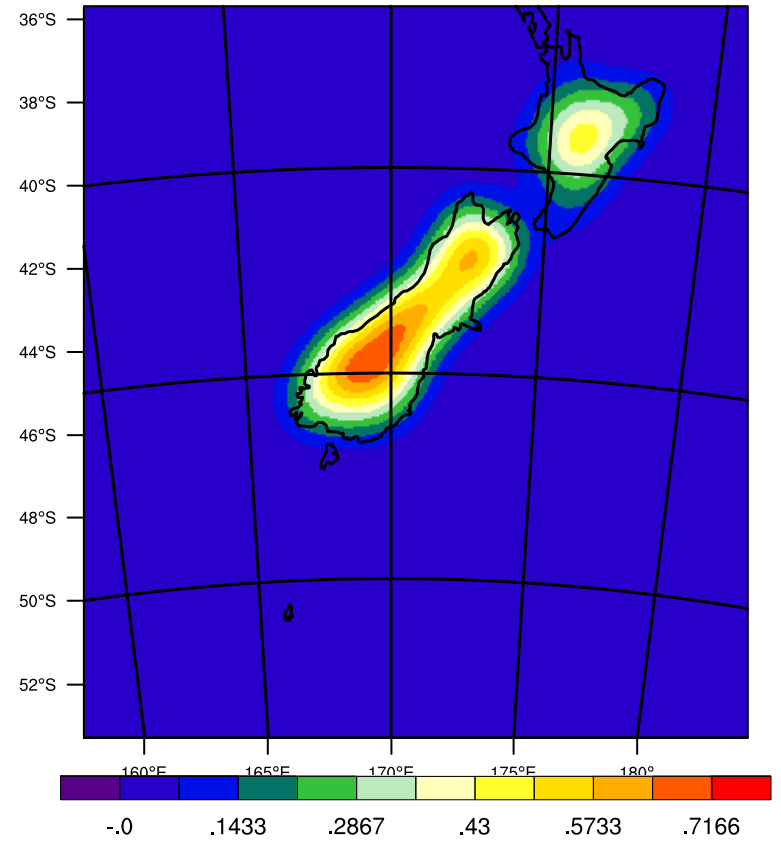
Init: 2011-07-10_06:00:00

Init: 2011-07-10_06:00:00

Terrain Height (m)

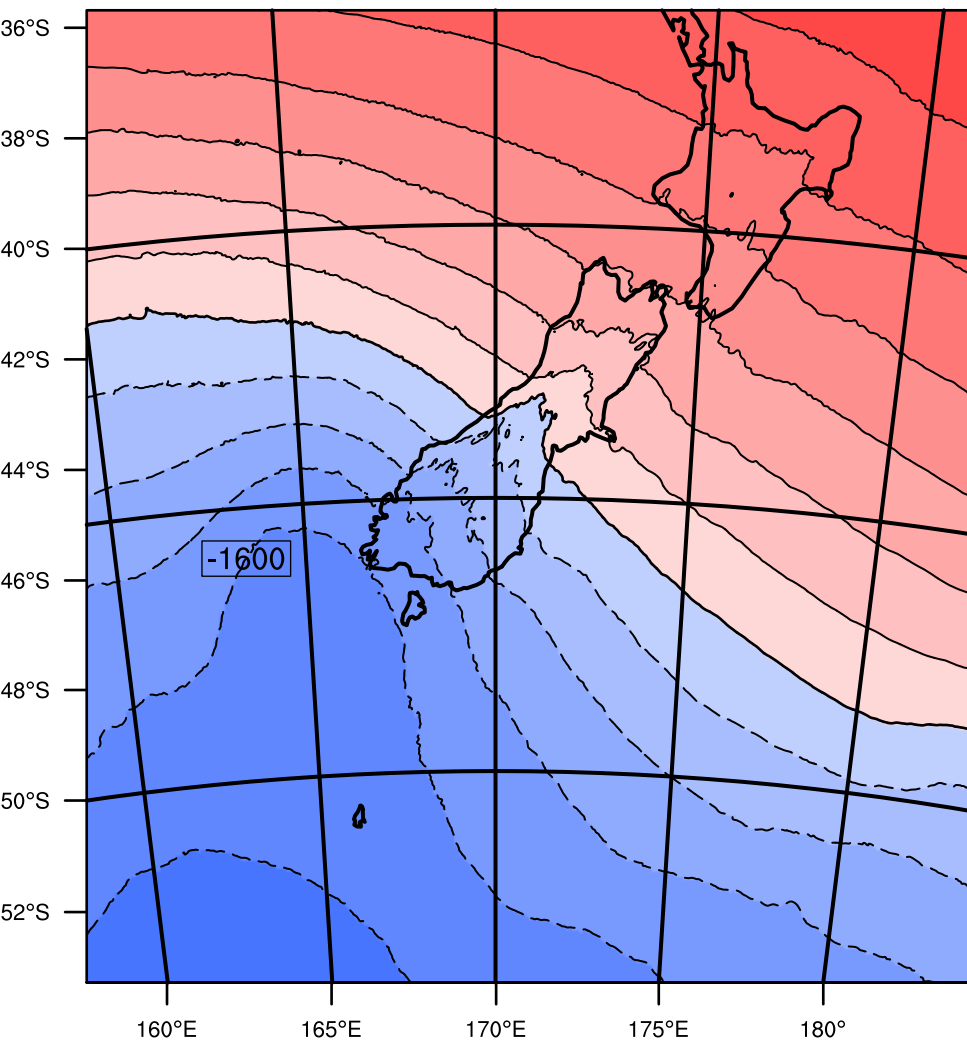


UnKnown



Demeaned 4 km pressure field

p-avg(p) (Pa) at 4 km
p-avg(p) (Pa) at 4 km

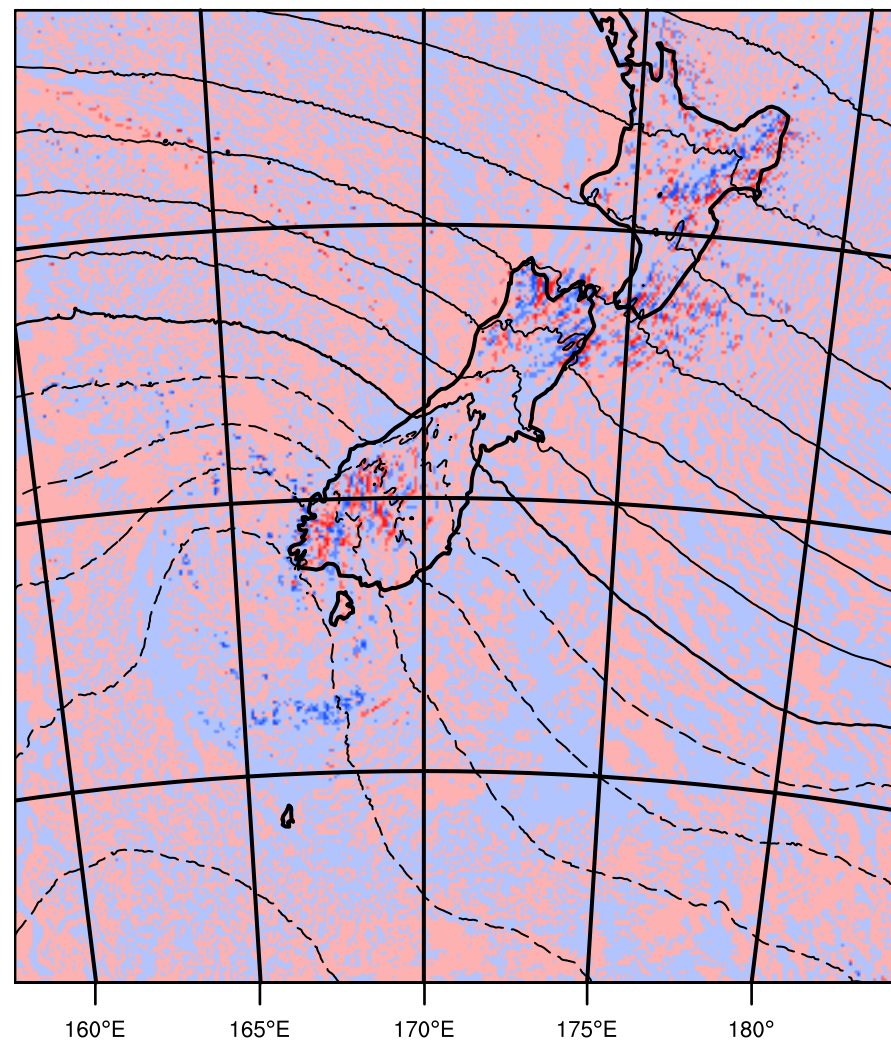


p-avg(p) Contours: -2800 to 3200 by 400
p-avg(p) (Pa)

-2400 -1600 -800 0 800 1600 2400 3200

EFz from demeaned p and w

(p-avg(p))*w (W m⁻²) at 4 km
p-avg(p) (Pa) at 4 km

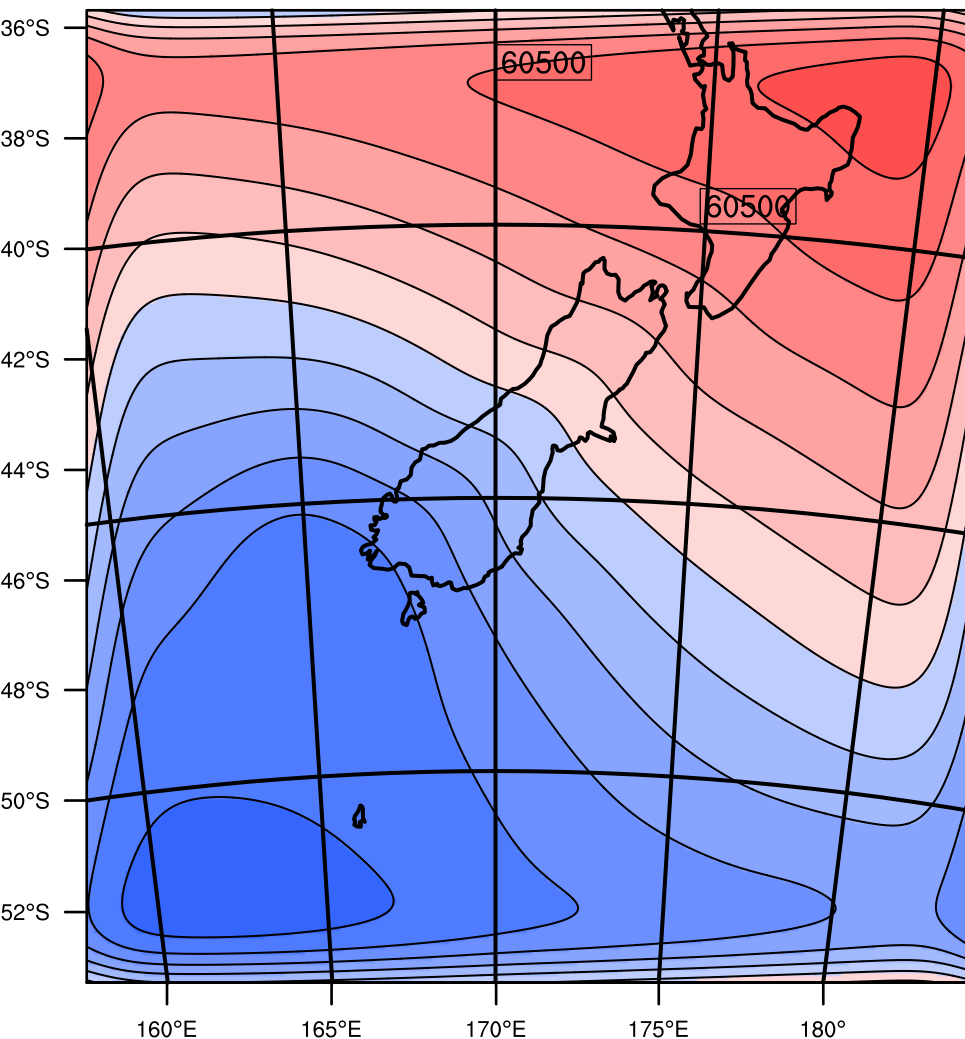


p-avg(p) Contours: -2800 to 3200 by 400
(p-avg(p))*w (W m⁻²)

-2000 -1000 0 1000 2000

Low passed pressure field

pressure LP @ 300 km (Pa) at 4 km
 pressure LP @ 300 km (Pa) at 4 km

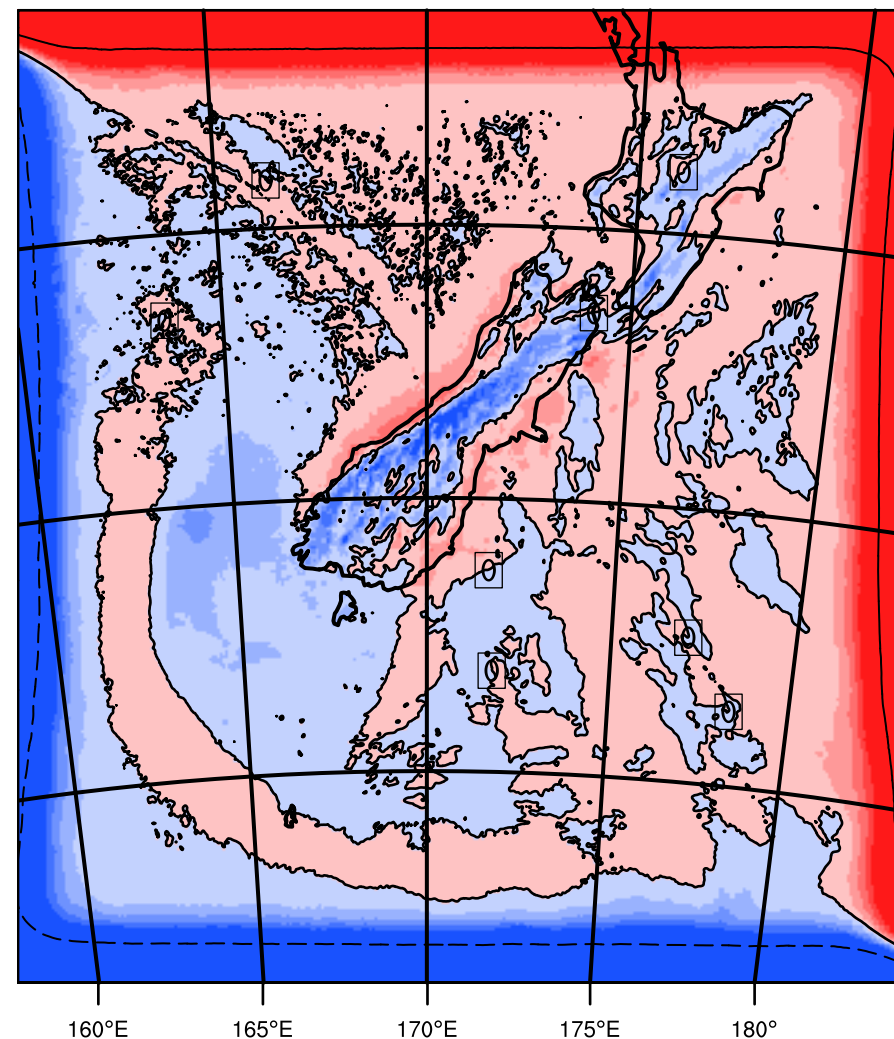


pressure LP @ 300 km Contours: 56500 to 61700 by 400
 pressure LP @ 300 km (Pa)

56500 57300 58100 58900 59700 60500 61300

High passed pressure field

pressure HP @ 300 km (Pa) at 4 km
 pressure HP @ 300 km (Pa) at 4 km



pressure HP @ 300 km Contours: -500 to 500 by 500
 pressure HP @ 300 km (Pa)

-200 -150 -100 -50 0 50 100 150 200