# DMS IN THE REMOTE, TROPICAL MBL

3<sup>RD</sup> PASE WORKSHOP December 2, 2009

Stephen Conley: Faloona, Blomquist, Miller, & Bandy UC Davis

#### **Scalar Budget Equation**

$$\frac{\partial c}{\partial t} + \nabla \cdot (\boldsymbol{u}c) = -L$$
$$I \qquad III \qquad III$$

I - Storage (rate of change of concentration.
 II - Flux Divergence
 III - Chemical source/sink

#### **Budget Simplifications**

• Flow is incompressible:  $\nabla \cdot u = 0$ 

Turbulence is horizontally homogeneous:  $\frac{\partial}{\partial x}(\overline{u'c'}) = 0; \frac{\partial}{\partial y}(\overline{v'c'}) = 0$ 

Simplified Budget Equation:

$$\frac{\partial C}{\partial t} = -U\frac{\partial C}{\partial x} - V\frac{\partial C}{\partial y} - \frac{\partial}{\partial z}(\overline{w'c'}) - L$$



### Estimating the budget terms

Actual DMS concentration is a complex function of position and time, i.e. c = f(x,y,z,t)

Expanding *f* in a Taylor series, retaining only first order terms leads to:

$$c = c_0 + \frac{\partial c}{\partial x}\Delta x + \frac{\partial c}{\partial y}\Delta y + \frac{\partial c}{\partial z}\Delta z + \frac{\partial c}{\partial t}\Delta t$$

### RF06 TBL2

RF06-TBL2



#### **RF06 DMS Concentration**



### **Horizontal Homogeneity?**

Chl-a Concentration (mg/m<sup>3</sup>)



First Two Weeks of PASE

Last Two Weeks of PASE

### Time Dependence



DMS Rose by 14 ppt/day - 70 ppt total!

#### Methodology

• Assumption:

$$c = \beta_0 + \beta_1 x + \beta_2 y + \beta_3 z + \beta_4 t + \beta_5 \cdot \sin(\frac{2\pi(t - \phi)}{24})$$

- Find the coefficients  $(\beta_n)$  such that the sum of the squares of the deviations between [DMS] calculated and actual is a minimum.
- Passed 1-second averages of *c,x,y,z,t* into Matlab's multiple linear regression solver to estimate the coefficients and the uncertainty in those coefficients.

### **DMS Fit Accuracy**



RF02: R<sup>2</sup> = 0.82

RF02: R<sup>2</sup> = 0.44

#### Vertical Flux Profile

Take the vertical derivative of budget equation:

$$\frac{\partial}{\partial z}\left(\frac{\partial C}{\partial t}\right) = -\frac{\partial}{\partial z}\left(U\frac{\partial C}{\partial x} - V\frac{\partial C}{\partial y}\right) - \frac{\partial^2}{\partial z^2}\left(\overline{w'c'}\right) - \frac{\partial L}{\partial z}$$

Switch the time and height derivatives and assume advection is constant with height. Given the small variability of OH in the MBL, assume the last term is zero also. This leads to the assumption of linearity in the flux profile, i.e.  $\frac{\partial^2}{\partial r^2}(\overline{w'c'}) = 0$ 

#### Vertical Flux Calculation

Flux = w'c'

w' = w - w

c' = c - c

What is the mean vertical wind?

## **Scales of Motion**



#### Cospectra of DMS and W



Cutoff corresponds to length scale of 20 km.

#### **RF03** Vertical DMS Flux Profile



## **DMS Budget**

RF	Obs ppt/hr	Div ppt/hr	Adv ppt/hr	Chem ppt/hr	ОН	Net Error
2	-5.2	0.9	0.5	-6.5	-7.0	0.4
3	-6.0	1.9	-0.8	-7.1	-7.0	0.4
5	-4.7	4.1	2.0	-10.9	-10.8	0.7
8	-5.6	0.8	-0.5	-5.8	-5.5	0.4
11	-5.9	1.3	2.3	-9.5	-6.4	1.4
12	-5.9	1.5	-1.1	-6.3	-7.0	0.5
14	-7.7	-0.8	-1.2	-5.7	-10.7	0.5
Avg	-5.8	1.4	0.2	-7.4	-7.8	0.6
6	-0.8	-0.2	1.8	-2.4	-2.2	0.6
13	-1.6	3.2	-2.5	-2.2	-3.4	0.7
Avg	-1.2	1.5	-0.4	-2.3	-3.4	0.6

### Dark Territory?

- Using the average daytime loss rate of ~6 ppt/ hour, an average day of photochemistry will consume ~50 ppt of DMS.
- Assuming vertical flux divergence (~1.4 ppt/ hour) is the only source of DMS, and that the buildup occurs over the 12 hours of darkness, we can only account for ~20 ppt during the entire night.
- Assuming that the system is in or near steady state, we are unable to account for more than half of the nightime buildup of DMS.

#### GRAPHIC OF DIURNAL OSCILLATION OF THE DMS ENTRAINMENT FLUX



## Nighttime Return



#### Nighttime Return

- From the 1994 Christmas Island data, the daily maximum occurs around 4:30AM.
- During PASE night flights, our first MBL leg occurred between 3AM and 4AM – right at the end of the increase.
- RF06 shows no significant nocturnal increase (stack #2 is 2 ppt higher than stack #1)
- RF13 does show a significant increase between stacks 1 and 2, nearly 10 ppt (4 ppt/hour)

## Nighttime Return



Stack #1 Flux Divergence = 0.6 ppt-m/s ÷ 550m × 3600 s/hr= 4 ppt/hr

### Kaimal Cospectra



Θ behaves much like the Kaimal curve

DMS & water vapor are shifted toward larger scales



LES simulation shows similar results, for  $\theta$  ( $\phi$ ~135) the peak occurs at smaller scales than it does for DMS ( $\phi$ ~60).

## The Mystery Flight!

Observed Decrease
Flux Divergence
Advection
Chemical Loss

-7.7 ppt/hr -0.8 ppt/hr -1.2 ppt/hr -5.7 ppt/hr

Predicted OH Loss

-10.7 ppt/hr



## **RF14 DMS Time Series**



## **RF14 Horizontal Gradient**



#### **RF14 Assumptions**

Flux profile not linear

Air mass far from steady state

Recently travelled over very productive water

Horizontal gradient difficult to measure

MBL height varied throughout the flight

#### **Flux Profile Evolution**

• Assuming  $w'c'_{z_i} = v_e \cdot ([c_{BuL} - c_{MBL}])$ 

Throughout the day, DMS in both layers is reduced by the same fraction.

 The magnitude of the difference decreases throughout the day.

 We expect the BuL flux to increase throughout the day.

## BuL - MBL



#### **DMS Flux into BuL**



DMS flux into BuL decreases throughout the day

## Conclusion

With the exception of RF14, the budget closure was excellent! More night flights!!

The budget closes without the need to invoke halogen chemistry.

During the flight hours of PASE, ~80% of the DMS entering the MBL from the surface is passed to the BuL. Need more night flights!!