

Cirrus cloud and ice supersaturation

Cirrus clouds have large but highly uncertain impacts on Earth's climate [Chen et al. 2000]. Modeling cirrus clouds is challenging because cloud microphysical processes occur at much smaller scales than climate model resolutions. Cirrus cloud formation requires supersaturation of the relative humidity with respect to ice (RHi). However, it has not been studied by using quasi-Eulerian aircraft observations how cirrus clouds initiate from ice supersaturated regions (ISSRs, regions where RHi > 100%), grow in size and eventually dissipate. Here we show the time evolution of cirrus clouds by analyzing the relationship between ice crystal regions (regions where ice crystal number density > $0.06 \ \text{#/cm}^3$) and ISSRs.



Deep convective clouds & chemistry experiment (DC3) campaign



The DC3 campaign includes 26 flights from May 18 - June 30, 2012, which took place over central plain of North America, investigating the impact of deep midlatitude continental convective clouds. The campaign provided large amount samplings around convective generated cirrus clouds, which is very helpful for analyzing ice crystal formation and evolution in convective regions. The analyses in this study were all restricted to temperature (T) \leq - 40 °C (~69.5 hrs), including ~11.2 hrs of ice supersaturation and ~21.1 hrs of ice crystal observations. The mean true air speed is 237 m/s.

GV instruments	Principal Investigator	Measurement	Accuracy
VCSEL hygrometer	Mark Zondlo, Princeton U.	Water vapor	6%
2-DC	Dave Rogers, NCAR	Ice particle number density (Nc) and mean diameter (Dc)	Measuremer 25-800 µm
CO_RAF	Teresa Campos, NCAR	CO	± 5%
O3_NCAR	Teresa Campos, NCAR	O3	± 9%

Definitions of ice crystal regions (ICRs) and ice supersaturated regions (ISSRs)

ISSRs: regions with spatially continuous ISS.

ICRs: regions with spatially continuous ice crystal distribution.

"With ice crystals" as where the ice crystal number density (Nc) \geq 0.06 #/cm³ during the 1 Hz measurements, while the remaining regions are considered to be clear-sky regions.

One ISSR+ICR sample: a set of spatially continuous ISSRs and ICRs. In total, **2529** ISSR+ICR samples were observed in DC3.

References:

Chen, T., W.B. Rossow and Y.C. Zhang (2000), Radiative effects of cloud-type variations. J. *Clim.*, 13, 264-286.

Diao, M., M.A. Zondlo, A.J. Heymsfield, D.C. Rogers and S.P. Beaton. Ice crystal evolution phases from *in situ* Eulerian observations. In preparation for Geophysical Research Letters. Fusina, F., P. Spichtinger and U. Lohmann (2007), Impact of ice supersaturated regions and thin cirrus on radiation in the midlatitudes. J. Geophys. Res. 112, D24S14, doi:10.1029/2007JD008449.

Rogers, R.R. and M.K. Yau (1989), A short course in cloud physics, Pergamon Press. Straka, J.M. Cloud and precipitation microphysics - principles and parameterizations, 1st ed., Cambridge University Press (2009).

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Ice crystal formation and evolution in DC3 campaign



Pha -se	Description	Spatial ratio M = sum(L _{ICR}) / L _{ISSR+ICR}
1	Clear-sky ISSRs	0
2	Nucleation	0 < M < 1
3	Early growth of ice crystals	0 < M < 1
4	Later growth of ice crystals	1
5	Evaporation/sedimentation	1





Phase 1: Clear-sky ISSRs phase: birthplace of cirrus clouds. : Nucleation phase: ICRs start to form inside the ISSRs. Phase 3: Early growth phase: ICRs and ISSRs are adjacent or they intersect each other. RHi decreases as ICR size expands.

4: Later growth phase: ICRs take over the space and ISSRs are imbedded in ICRs. The whole ICR+ISSR is no longer fully supersaturated. Phase 5: Ice crystals evaporate and sediment. No more ISS. RHi decreases as ICR size shrinks. Note the wide distribution of RHi (from 100% to 10%) for aged ICRs.

Precision

≤ 1%

nt range:

± 5%

0.8 ppbv



322 88 152 0 < N < 1 266 $0 < N \leq 1$ 1701



similar values

0.001 0.01 0.1





[1] DC3 campaign shows more ICRs with ISS buried inside (Phase 4), while in START08 and HIPPO Global campaigns, ISS disappears faster once ice crystals are formed, suggesting in DC3 ice crystals do not consume water vapor over saturation as fast as those in START08 and HIPPO.

[2] ICR/ISSR ratio peaks at higher value in DC3 than the other two campaigns, suggesting that ice crystals grow relatively fast in DC3 so that most ICRs are larger than ISSRs, which indicates that the dominance of Phase 4 in DC3 is contributed by continued strong uplift which maintains the ISS inside ICRs instead of inefficient depositional growth.

Implications to understanding ice crystal evolution

1. We provided a new and simple method to distinguish three phases of ice cloud evolution by using in-situ and quasi-Eulerian sampling of two common parameters: 1) RHi 2) presence/absence of ice crystals. 2. We demonstrated the importance of separating out various evolution phases of ice crystals in Eulerian view observation, since they have different properties of RHi, Nc and Dc.

3. We compared ice crystal evolution between various geographical location and meteorological background. Our finding improves the understanding of the relative lifetime of ice crystal evolution. 4. Our result facilitates the comparison between in situ Eulerian observation and Lagrangian view cloud simulation/ parameterization.



[1] Mean diameter of ice crystals merges into a constant value as ice crystals grow, which

[2] Number density of ice crystals continues to increase throughout the ICR evolution. The increasing Nc agrees with previous simulations, where new ice crystals continue to form as the air parcel continues to be uplifted [Spichtinger and Gierens, 2009].

Ice crystal lifetime phases and comparisons with other campaigns (START08 & HIPPO1-5)